Two-dimensional turbulence: a review

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3d equations of motion

The equation of motion for an incompressible 3d flow are



 $\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} = -\frac{\nabla p}{\rho} + v \nabla^2 \mathbf{u}$

 $(\nabla \cdot \mathbf{u} = \mathbf{0}, \mathbf{v} \text{ viscosity})$

George G. Stokes, 1845

Claude Navier, 1827

3d NS equations conserves kinetic energy $E = \frac{1}{2} \int |\mathbf{u}|^2 d^3 x$ in the inviscid limit.

The viscous energy balance is

$$\frac{dE}{dt} = -2vZ$$

enstrophy $Z = \frac{1}{2} \int \left| \nabla \times \mathbf{u} \right|^2 d^3 x$

Navier-Stokes equations and turbulence

The flow becomes turbulent when

$$\mathsf{Re} = \frac{\left[\mathbf{u} \cdot \nabla \mathbf{u}\right]}{\left[\nu \nabla^2 \mathbf{u}\right]}; \frac{UL}{\nu}? 1$$

In this limit (i.e. $v \rightarrow 0$) the energy dissipation rate remains finite:

 $\varepsilon = \lim_{v \to 0} 2vZ > 0$ dissipative anomaly

Typical Re numbers laboratory: Re $\approx 10^4$ atmosphere: Re $\approx 10^7$



Re







Fig. 5.11. Variation of drag coefficient with Reynolds number for circular cylinders.

Ekman-Navier-Stokes equations

3D Navier Stokes



Two-dimensional turbulence and geophysical flows

2D Navier-Stokes equation are a simple model for large scale motion of atmosphere and oceans: thin layers of fluid in which stratification and rotation supress vertical motions.



Energy/enstrophy balance for 2d NS

2d Navier-Stokes equations have two inviscid quadratic invariants:

Energy $E = \frac{1}{2} \int |\mathbf{u}|^2 d^2 x = \int E(k) dk$ Enstrophy $Z = \frac{1}{2} \int \omega^2 d^2 x = \int k^2 E(k) dk$

Energy/enstrophy balance in viscous flow:

$$\frac{dE}{dt} = -2\nu Z \qquad \qquad \frac{dZ}{dt} = -2\nu P$$

$$P = \frac{1}{2} \int \left| \nabla \times \omega \right|^2 d^2 x \ge 0$$

palinstrophy

E(k): energy spectrum

In fully developed turbulence limit, Re=UL/v -> ∞ (i.e. v->0):

$$\lim_{v\to 0}\frac{dE}{dt}=0$$

(because $dZ/dt \le 0$ and $Z(t) \le Z(0)$)

no dissipative anomaly for energy in 2d: no energy cascade to small scales !

The double cascade

In the limit of Re-> ∞ 2d turbulence displays an enstrophy cascade to small scales at a rate ζ . Energy flows to large scales with rate ε generating the inverse cascade.

Two inertial range of scales: •energy inertial range $1/L < k < k_F$ (with constant ε) •enstrophy inertial range $k_F < k < k_d$ (with constant ζ)

(Kraichnan 1967, Leith 1968, Batchelor 1969)





log(k)

Two power-law self similar spectra in the inertial ranges.

log(E(k))

The double cascade scenario is typical of 2d flows, e.g. plasmas and geophysical flows.

Exact results

Following the derivation obtained by Kolmogorov for 3d turbulence (Kolmogorov 4/5 law) is it possible to obtain for 2d cascades two exact results:

inverse energy cascade:

$$S_3(r) = \left\langle \left(\delta_r u \right)^3 \right\rangle = \frac{3}{2} \varepsilon r$$

direct enstrophy cascade:

$$\left\langle \delta_{r} u \left(\delta_{r} \omega \right)^{2} \right\rangle = -2\zeta r$$

Kolmogorov's 4/5 law (1941)



Dissipation of energy in the locally isotropic turbulence[†]

By A. N. Kolmogorov

In my note (Kolmogorov $1941\,a)$ I defined the notion of local isotropy and introduced the quantities

$$\begin{array}{l} B_{dd}(r) = \overline{[u_d(M') - u_d(M)]^2}, \\ B_{nn}(r) = \overline{[u_n(M') - u_n(M)]^2}, \end{array} \tag{1}$$

where r denotes the distance between the points M and M', $u_d(M)$ and $u_d(M')$ are the velocity components in the direction $\overline{MM'}$ at the points M and M', and $u_n(M)$ and $u_n(M')$ are the velocity components at the points M and M' in some direction, perpendicular to MM'.

In the sequel we shall need the third moments

$$B_{ddd}(r) = \overline{[u_d(M') - u_d(M)]^3}.$$
 (2)

For the locally isotropic turbulence in incompressible fluid we have the equation

$$4E + \left(\frac{\mathrm{d}B_{ddd}}{\mathrm{d}r} + \frac{4}{r}B_{ddd}\right) = 6\nu \left(\frac{\mathrm{d}^2B_{dd}}{\mathrm{d}r^2} + \frac{4}{r}\frac{\mathrm{d}B_{dd}}{\mathrm{d}r}\right) \tag{3}$$

similar to the known equation of Kármán for the isotropic turbulence in the sense of Taylor. Herein \overline{E} denotes the mean dissipation of energy in the unit of time per unit of mass. The equation (3) may be rewritten in the form

$$\frac{\mathrm{d}}{\mathrm{d}r} + \frac{4}{r} \left(6\nu \frac{\mathrm{d}B_{dd}}{\mathrm{d}r} - B_{ddd} \right) = 4\overline{E}, \qquad (4)$$

and, in virtue of the condition $(d/dr)B_{dd}(0) = B_{ddd}(0) = 0$, yields

$$6\nu dB_{dd}/dr - B_{ddd} = \frac{4}{5}\overline{E}r. \qquad (5)$$

For small r we have, as is known,

i.e.

$$B_{dd} \sim \frac{1}{15} \overline{E} r^2 / \nu, \qquad (6)$$

$$6\nu dB_{aa}/dr \sim \frac{4}{\xi} Er$$
.

Thus, the second term on the left-hand side of (5) is for small r infinitesimal in comparison with the first. For large r, on the contrary, the first term may be neglected in comparison with the second, i.e. we may assume that

$$B_{ddd} \sim -\frac{4}{5}\overline{E}r.$$
 (7)

It is natural to assume that for large r the ratio

$$S = B_{ddd} : B_{dd}^3, \tag{8}$$

† First published in Russian in Dokl. Akad. Nauk SSSR (1941), 32(1). Paper received 30 April 1941. This translation by V. Levin, reprinted here with emendations by the editors of this volume.

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Proc. R. Soc. Lond. A (1991) 434, 15–17 Printed in Great Britain





Atmospheric flows:

Mesoscale wind variability (radar and balloon): k^{-5/3} K.S. Gage, J.Atmos.Sciences **36** (1979) QuickTime[™] and a YUV420 codec decompressor are needed to see this picture.

GASP aircraft dataset: k^{-5/3} for wavelenghts 3-300 km Nastrom, Gage, Jasperson, Nature 310 (1984)



Laboratory experiments

Soap films

(Y. Couder, W. Goldburg, H. Kellay, M.A. Rutgers, M. Rivera, R.E. Ecke)

interferometry, LDV, PIV

Thin layer of electrolyte driven by Lorentz force

(P. Tabeling, J. Gollub, A. Cenedese)

PIV





Direct numerical simulations

Integration of 2d-NS equations for vorticity scalar field $\omega {=} {-} \nabla^2 \psi$

$$\frac{\partial \omega}{\partial t} + \mathbf{u} \cdot \nabla \omega = v \nabla^2 \omega + f$$

stream function ψ with $\mathbf{u} = (\partial_{y}\psi, -\partial_{x}\psi)$

Homogeneous, isotropic turbulence: square box with periodic boundary conditions

Pseudo-spectral code

* Fourier decomposition: derivatives in Fourier spaces
* Convolutions (products): back in physical space with FT

Extensive use of Fast Fourier Transform Computationally much faster than 3d simulations

A brief history of DNS

DNS of 2d turbulence followed and explosive trend in the early stage.

From Frisch & Sulem (Phys. Fluids 27 1984) at 256² to Borue (Phys. Rev. Lett. 71, 1993) at 4096² the resolution grows following an exponential law:

N = 267*2(+-1984)/2.8



Late stage: saturation at "reasonable" resolution (present simulations at 2048² on PC)

High resolution DNS

(G. Boffetta and A. Celani, 2005)

Set of simulations with resolutions up to 16384² with a parallel pseudospectral code on IBM-SP4 at Cineca.

Ν	ν	α	r _F	η	3	ζ
2048	2×10 ⁻⁵	0.015	0.01	2.4×10 ⁻³	2.1×10 ⁻³	38.0
4096	5×10-6	0.024	0.01	1.2×10 ⁻³	3.2×10 ⁻³	36.1
8192	2×10 ⁻⁶	0.025	0.01	7.8×10 ⁻⁴	3.6×10 ⁻³	35.3
16384	1×10-6	0.0	0.01	5.5×10 ⁻⁴	3.6×10 ⁻³	37.6

Simultaneously observation of direct and inverse cascade



Energy spectra



a bracificanthabh) nasangmaabh CiadhiVintenaraC

Direct cascade

Early numerical simulations failed to reproduce k⁻³ spectrum Steeper spectrum (k⁻⁴-k⁻⁵) associate with the presence of strong large scale vortices Decaying turbulence shows long term memory of initial conditions (no universality) (Benzi, McWilliams)

Recent simulations more in agreement with Kraichnan prediction (logarithmic corrections?) S.Chen, R. Ecke, G. Eyink, X. Wang, Z. Xiao, PRL 91, (2003) E.Lindborg, K.Alvelius, Phys. Fluids 12 945 (2000) C.Pasquero, G.Falkovich, PRE 65 056305 (2002) QuickTime[™] and a TIFF (Uncompressed) decompressor are needed to see this picture.

Direct cascade with linear friction: when friction is sufficiently strong small scale vorticity has the same statistics of a passive scalar with finite lifetime (linear problem)



Mean field computation for passive scalar with finite lifetime

S. Corrsin, JFM 11, 407 (1961)

Stretching: incompressible smooth velocity generates small scale features at exponential rate (Lyapunov):

$$r = L e^{-\lambda t}$$

Linear damping: fluctuations decay at exponential rate

$$\theta(\mathbf{x}(\mathbf{1}),\mathbf{1}) = \theta(\mathbf{x}(\mathbf{0}),\mathbf{0}) \mathbf{e}^{-\alpha}$$



Combining the two exponential behavior, one has $\delta\theta(r,t) = \delta\theta(L,0) \left(\frac{r}{L}\right)^{\alpha/\lambda}$ and thus, in stationary conditions: and thus, in stationary conditions: $\left\langle \left(\delta \theta(\mathbf{r}, \mathbf{t}) \right)^{2p} \right\rangle \approx \mathbf{r}^{2p\alpha/\lambda}$

 α -dependent structure function exponents (and spectral index)

Fluctuations of Lyapunov exponent: intermittency corrections

M. Chertkov, Phys. Fluids 10, 3017 (1998)



 $T_{L}(r)$ is the time that a couple of particles at distance r takes to reach a separation L_{F} backward in time.

In a two-dimensional incompressible smooth flow particles separate exponentially fast

$$\Box > T_{L}(r) = \frac{1}{\gamma} \ln(L/r)$$

 $L = r e^{\gamma t} \quad)$

Link between exit-time $T_L(r)$ and finite time Lyapunov exponent γ

Statistics of vorticity

•Vorticity structure functions:

 $S_{p}^{\omega}(\mathbf{r}) \equiv \left\langle \left| \delta \omega (\mathbf{r}) \right|^{p} \right\rangle; \left\langle \Omega^{p} \right\rangle_{f} \left\langle \left(\frac{\mathbf{r}}{L} \right)^{\frac{p\alpha}{\gamma}} \right\rangle$

•Distribution of finite time Lyapunov exponent γ (at long times):

 $G(\gamma)$

$$P(\gamma, t) \approx t^{1/2} e^{-G(\gamma)t}$$



 $G(\gamma)$ (Cramér function) has a quadratic minimum at $\gamma = \lambda$ (Lyapunov exponent)

Prediction for structure functions

$$\mathcal{P}) \approx \left\langle \Omega^{p} \right\rangle \int d\gamma \, \left(\frac{\mathbf{r}}{\mathsf{L}} \right)^{\frac{pa+b(\gamma)}{\gamma}}$$

$$\left(\frac{r}{L}\right)$$

 $\zeta(p)$

$$f(p) = \min_{\gamma} \left\{ p\alpha + G(\gamma) \right\} \gamma$$

 $S^{\omega}_{p}(r)$

* scaling exponents depend on friction coefficient $\boldsymbol{\alpha}$

* intermittency: nonlinear $\zeta(p)$ for generic G for quadratic $G(\gamma)=(\gamma-\lambda)^2/2\mu$ one has the explicit expression

$$\zeta(\boldsymbol{p}) = \frac{1}{\mu} \left[\sqrt{\lambda^2 + 2\boldsymbol{p}\alpha\mu - \lambda} \right]$$

Correction to spectral slope $E(k) \approx k^{-3-\zeta(2)}$ velocity is smooth

and Charl



Numerical results: effects of friction on spectrum



Energy spectrum always steeper than k⁻³: smooth velocity field

 $E(k) \approx k^{-3-\zeta(2)}$

Correction to spectral slope depends on friction intensity



Vorticity fields

QuickTime™ and a Sorenson Video 3 decompressor are needed to see this picture. QuickTime™ and a Sorenson Video 3 decompressor are needed to see this picture.

α=0.15

α=0.05

Comparing vorticity and passive scalar





Intermittency

Probability density functions $p(\delta \omega(r))$ not self-similar in r

Identical mechanism for vorticity and passive scalar: fluctuations of γ

Intermittency: anomalous scaling of vorticty structure functions

Prediction in terms of Cramer function

$$\zeta(p) = \min_{\gamma} \left\{ p\alpha + G(\gamma) \right\} \gamma$$



direct computation from SF prediction from Cramer function Experimental study of direct cascade with friction
G. Boffetta, A. Cenedese, S. Espa, S. Musacchio *Europhysics Letters* 71, 590 (2005).
Correction to the spectral slope observed
in experiments with electrolyte cell
α≈v/h²



Magnets arrangement Videocamera

Inverse cascade

* Kraichnan-Kolmogorov spectrum k^{-5/3} well observed
* absence of intermittency
* quasi-Gaussian statistics

> QuickTime™ and a TIFF (LZW) decompressor are needed to see this picture.

Early studies on inverse cascade

Thin layer of mercury with electromagnetic forcing J.Sommeria, JFM **170**, 139 (1986)





Pseudo-spectral DNS at resolution 256² U.Frisch, P.L. Sulem, Phys. Fluids **27**, 1921 (1984)

More recent experiments ...



Electrolyte cell J. Paret, P. Tabeling, PRL **79** 4162 (1997)

... and numerical simulations



L.M.Smith, V.Yakhot, PRL 71 352 (1993)



 10^{2} 100 10 10⁰ ^{/3} ^{5/3}E(k) 0.1 1 $10^1 10^2 10^3$ 10⁻² E(K) 0.2 П(k) 10⁻⁴ 0.1 -0.1 10⁻⁶ .-0.2 10² 10³ 10¹ 10 100 1000 1

G.Boffetta, A.Celani, M.Vergassola, PRE 61 R29 (2000)

much wider inverse cascade range than experiments!

Kolmogorov law in inverse cascade (constant energy flux)

$$\left\langle \left(\delta u_{//}(r) \right)^3 \right\rangle = \frac{3}{2} \varepsilon r$$

between forcing scale r_{F} and friction scale $r_{fr}\text{=}\epsilon^{1/2}\alpha^{-3/2}$



Energy spectrum and spectral flux $E(k) = C\varepsilon^{2/3}k^{-5/3}$ $C = 6.0 \pm 0.4$ $\left\langle \left(\delta u_{//}(r) \right)^2 \right\rangle; \ 12.9(\varepsilon r)^{2/3}$



Higher order structure functions

$$S_{p}(r) = \left\langle \left(\delta u_{//}(r) \right)^{p} \right\rangle = C_{p} \left(\varepsilon r \right)^{p/2}$$

compatible with Kolmogorov scaling no intermittency



Skewness is small (smaller than in 3d):

 $S = \frac{\left\langle \delta u(r)^3 \right\rangle}{\left\langle \delta u(r)^2 \right\rangle^{3/2}} \approx 0.03$

pdf of velocity fluctuations not far from Gaussian (compared with 3d)



Deviations from Gaussian pdf

Antisymmetric part of the pdf:

 $p_A(u)=p(u)-p(-u)$ is a measure of the deviation from Gaussian distribution

Antisymmetric part of longitudinal velocity increment pdf at separations r=0.05, 0.075, 0.1 (rescaling)

Inset: antisymmetric part (lower) compared with symmetric part (upper) at r=0.1



Core of the distribution not far from Gaussian, but very large deviations for large fluctuations δu (to be included in closures?)

Experimental study of two-dimensional enstrophy cascade G. Boffetta, A. Cenedese, S. Espa and S. Musacchio

Europhysics Letters 71, 590 (2005).

Intermittency in two-dimensional Ekman-Navier-Stokes turbulence G. Boffetta, A. Celani, S. Musacchio and M. Vergassola, *Physical Review E* 66, 026304 (2002).

Closure of two dimensional turbulence: the role of pressure gradients G. Boffetta, M. Cencini and J. Davoudi, *Physical Review E* **66**, 017301 (2002).

Inverse energy cascade in two-dimensional turbulence: Deviations from Gaussian behavior

G. Boffetta, A. Celani and M. Vergassola, *Physical Review E* 61, R29 (2000).

http://www.ph.unito.it/~boffetta

Statistical description

<...> ensemble averaging

Two point correlations
$$C_{\alpha\beta}(\mathbf{r}) = \langle u_{\alpha}(\mathbf{x})u_{\beta}(\mathbf{x}+\mathbf{r}) \rangle = \sum_{k} e^{-k\mathbf{r}} \hat{C}_{\alpha\beta}(\mathbf{k})$$

Structure functions: $S_{\alpha\beta}(\mathbf{r}) = 2C_{\alpha\beta}(\mathbf{0}) - 2C_{\alpha\beta}(\mathbf{r}) = \langle \delta_{r}u_{\alpha}\delta_{r}u_{\beta} \rangle$

Using incompressibility and isotropy:

$$\hat{\mathcal{C}}_{\alpha\beta}(\mathbf{k}) = \mathcal{B}(\mathbf{k}) \left[\delta_{\alpha\beta} - \frac{\mathbf{k}_{\alpha} \mathbf{k}_{\beta}}{\mathbf{k}^{2}} \right]$$

energy spectrum:

$$E(k) \equiv 2\pi k^2 B(k)$$

Power-law energy spectrum

In scaling invariant inertial ranges one expects $E(k) \approx k^{-\gamma}$

for $1 \le \gamma \le 3$ (locality) one has

$$S_{\alpha\alpha}(r); r^{\gamma}$$

Higher order structure functions

$$S_{p}(r) \equiv \left\langle \left(\delta_{r} \mathbf{u} \cdot \hat{\mathbf{r}} \right)^{p} \right\rangle$$

$$\delta_r \mathbf{u} \equiv \mathbf{u}(\mathbf{x} + \mathbf{r}) - \mathbf{u}(\mathbf{x})$$